

Perturbations in spherically symmetric solutions

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- *On Horizons and Wormholes in k-Essence Theories*
(Bronnikov et al., 2016; [arXiv:1511.08036](#))
- *Static, spherically symmetric solutions with a scalar field in Rastall gravity* (Bronnikov et al., 2016; [arXiv:1606.06242](#))
- *Duality between k-essence and Rastall gravity*
(Bronnikov et al., 2017; [arXiv:1701.06662](#))

The k-essence theories can be defined by the following most general Lagrangian:

$$\mathcal{L} = \sqrt{-g}[R - F(X, \phi)]$$

Starting with the model

$$F(X) = F_0 X^n - 2\Lambda, \quad F_0, \Lambda = \text{const},$$

where

$$X = \eta \phi_{;\rho} \phi^{;\rho},$$

where $\eta = \pm 1$ is set to avoid ill-defined power-law.

Now, consider a general static, spherically symmetric metric

$$ds^2 = e^{2\gamma} dt^2 - e^{2\alpha} du^2 - e^{2\beta} (d\theta^2 + \sin^2 \theta d\phi^2),$$

where γ , α and β functions of u only.

- Solution 1: $n = \frac{1}{3}$ e $\Lambda = 0$

$$ds^2 = \frac{B(x)}{k^2 x} dt^2 - \frac{k^2 x}{B(x)} dx^2 - \frac{1}{k^2 x} d\Omega^2, \quad B(x) = B_0 - \frac{1}{2} k^2 x^4,$$

B_0 and k are constants.

- Solution 2: $n = \frac{1}{2}$

$$ds^2 = \frac{(9b^2 - x^2)^2}{9b^4} dt^2 - \frac{9b^4}{(9b^2 - x^2)^2} dx^2 - \frac{3b^4}{9b^2 - x^2} d\Omega^2,$$

with $b = \sqrt{\frac{3}{\Lambda}}$

Rastall gravity field equations

$$R^\nu_\mu - \frac{\lambda}{2} \delta^\nu_\mu R = -T^\nu_\mu,$$

$$\nabla_\nu T^\nu_\mu = \frac{\lambda - 1}{2} \partial_\mu R,$$

where RG is recovered with $\lambda = 1$.

$$G^\nu_\mu = - \left\{ T^\nu_\mu - \frac{b-1}{2} \delta^\nu_\mu T \right\} \equiv -\tilde{T}^\nu_\mu,$$

$$\nabla_\nu T^\nu_\mu = \frac{b-1}{2} \partial_\mu T, \quad b := \frac{3\lambda - 2}{2\lambda - 1}, \quad T = T^\alpha_\alpha.$$

In this parameterization, RG is recovered with $b = 1$.

- Solution for $b = -1$:

$$ds^2 = A(x)dt^2 - A(x)^{-1}dx^2 - \sqrt{\frac{3}{2C^2}}\frac{1}{x}d\Omega^2,$$

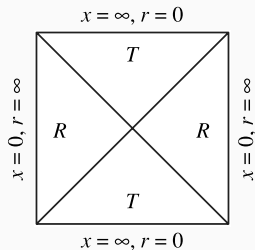
with $A = \frac{K}{x} - \frac{C}{\sqrt{6}}x^3$.

- Solution for $b = 0$:

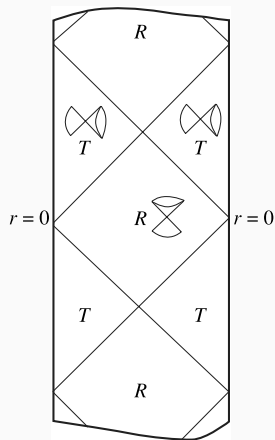
$$ds^2 = \frac{9b^4}{\cosh^4(bu)}dt^2 - du^2 - \frac{\cosh^2(bu)}{3b^2}d\Omega^2,$$

with $b = \sqrt{\frac{\Lambda}{3}}$

Carter-Penrose diagram



Solution for $n = 1/3$ / $b = -1$.



Solution for $n = 1/2$ / $b = 0$.

Adopting, now, a spherically symmetric metric

$$ds^2 = e^{2\gamma(u,t)} dt^2 - e^{2\alpha(u,t)} du^2 - e^{2\beta(u,t)} d\Omega^2,$$

Considering the perturbed quantities

$$\alpha(u, t) = \alpha(u) + \delta\alpha(u, t),$$

$$\beta(u, t) = \beta(u) + \delta\beta(u, t),$$

$$\gamma(u, t) = \gamma(u) + \delta\gamma(u, t),$$

$$\phi(u, t) = \phi(u) + \delta\phi(u, t)$$

$$F_X e^{\alpha+2\beta-\gamma} \ddot{\phi} - \left(F_X e^{-\alpha+2\beta+\gamma} \phi' \right)' = 0,$$

$$-e^{2\alpha-\beta} + \beta'(\beta' + 2\gamma') = \left(\frac{1}{2}F - XF_X \right) e^{2\alpha},$$

$$-e^{2\alpha-2\beta} + 2\beta'' + 3\beta'^2 - 2\alpha'\beta' = \frac{1}{2}F e^{2\alpha},$$

$$e^{2\alpha-2\beta} + e^{2\alpha-2\gamma} \ddot{\beta} - \beta''\beta'(\gamma' - \alpha' + 2\beta') = -\frac{1}{2}e^{2\alpha}(F - XF_X),$$

$$\dot{\beta}' + \dot{\beta}\beta' - \dot{\alpha}\beta' - \dot{\beta}\gamma' = \frac{1}{2}F_X \dot{\phi}\phi'.$$

and perturbed quantities for X and $F(X, \phi)$ are given by

$$\delta X = 2e^{-2\alpha}(\phi'\delta\phi' - \phi'^2\delta\alpha), \quad F(X, \phi) = F(X) + F_X\delta X + F_\phi\delta\phi,$$

$$F_X(X, \phi) = F_X(X) + F_{XX}\delta X + F_{X\phi}\delta\phi.$$

Perturbed scalar field equation

$$-e^{2\alpha-2\gamma}\delta\ddot{\phi} + \delta\phi'' + \phi'(2\delta\beta' + \delta\gamma' - \delta\alpha')$$

$$(2\beta' + \gamma' - \alpha')\delta\phi' + \frac{F'_X}{F_X}\delta\phi' + \phi'\delta\left(\frac{F'_X}{F_X}\right) = 0,$$

Adopting $\delta\beta = 0$

$$-e^{2\alpha-2\gamma}\delta\ddot{\phi} + \delta\phi'' + \phi'(\delta\gamma' - \delta\alpha')$$

$$+(2\beta' + \gamma' - \alpha')\delta\phi' + \frac{F'_X}{F_X}\delta\phi' + \phi'\delta\left(\frac{F'_X}{F_X}\right) = 0,$$

or

$$\begin{aligned} -e^{2\alpha-2\gamma}\delta\ddot{\phi} + (2n-1)\delta\phi'' + \delta\phi'[2\beta' + \gamma' - (2n-1)\alpha'] \\ + \phi'[\delta\gamma' - (2n-1)\delta\alpha'] = 0. \end{aligned}$$

Schrödinger-like equation

We can rewrite $\delta\phi$ equation

$$-e^{2\alpha-2\gamma}\delta\ddot{\phi} + (2n-1)\delta\phi'' + \delta\phi'[\sigma' - 2(n-1)\alpha'] - \frac{n^2}{\beta'^2}(e^{-2\beta} - \Lambda)e^{4\alpha}F_0X^n\delta\phi = 0.$$

Using transformations

$$\frac{du}{dz} = e^{\gamma-\alpha}, \quad \delta\phi = \Psi(z)e^{i\omega t}, \quad \Psi(z) = f(z)\psi(z),$$

with

$$f(z) = \exp\left[-\frac{\beta + (1-n)\gamma}{2n-1}\right].$$

We found the Schrödinger-like equation

$$(2n-1)\frac{d^2\psi}{dz^2} + [\omega^2 - V(x)]\psi(z) = 0,$$

Solution for $n = 1/3$

Replacing the solution for $n = 1/3$, we obtain

$$\psi_{zz} - [3\omega^2 + V(z)]\psi = 0,$$

$$\text{with } V(x) = -F_1 B(x) + \frac{5B_0^2}{4k^4 x^4} + \frac{31B_0}{4k^2} - \frac{3x^4}{16}.$$

The potential, near $x = 0$

$$z \propto x^2 \text{ as } x \rightarrow 0, \quad V(z) \approx \frac{5B_0}{4k^4 x^4} \propto \frac{1}{z^2}.$$

On the other hand, near $x = x_h$,

$$z \propto -\ln(x_h - x) \rightarrow \infty, \quad V(z) \rightarrow -\frac{8B_0}{k^2}.$$

What about ψ ?

$$\psi \sim e^z \text{ as } z \rightarrow \infty, \quad \psi \sim z^{-1/4} \text{ as } z \rightarrow 0.$$

Solution for $n = 1/3$

- We see that these requirements are much milder than could be imposed on a quantum-mechanical wave function.
- Therefore, ω^2 can take negative values arbitrarily large in absolute value, $e^{|\omega|t}$, and thus our static configuration is catastrophically unstable.

Also, the expression of sound speed for k-essence is

$$c_s^2 = \frac{F_X}{F_X + 2XF_{XX}},$$

or, in our model,

$$c_s^2 = \frac{1}{2n - 1},$$

where c_s is imaginary for $n < 1/2$.

$$- e^{2(\alpha-\gamma)}(\ddot{\alpha} + 2\ddot{\beta}) + \gamma'' + \gamma'(\gamma' - \alpha' + 2\beta') =$$

$$- \frac{\epsilon}{2}(1-a)\phi'^2 - (3-2a)e^{2\alpha}V,$$

$$- e^{2(\alpha-\gamma)}\ddot{\alpha} + \gamma'' + 2\beta'' - \alpha'(\gamma' + 2\beta') + \gamma'^2 + 2\beta'^2 =$$

$$- \frac{\epsilon}{2}(3-a)\phi'^2 - (3-2a)e^{2\alpha}V,$$

$$- e^{2(\alpha-\gamma)}\ddot{\beta} + \beta'' + \beta'(\gamma' - \alpha' + 2\beta') - e^{2(\alpha-\beta)} =$$

$$- \frac{\epsilon}{2}(1-a)\phi'^2 - (3-2a)e^{2\alpha}V,$$

$$\dot{\beta}' + (\beta' - \gamma')\dot{\beta} - \beta'\dot{\alpha} = -\frac{\epsilon}{2}\phi'\dot{\phi},$$

$$a\phi'' + [\gamma' - a\alpha' + 2\beta']\phi' - e^{2(\alpha-\gamma)}\ddot{\phi} = \epsilon(3-2a)e^{2\alpha}V_{\phi}$$

$$\begin{aligned}
 & -e^{2(\alpha-\gamma)}(\delta\ddot{\alpha} + 2\ddot{\delta}\beta) + \delta\gamma'' + \delta\gamma'(\gamma' - \alpha' + 2\beta') + \gamma'(\delta\gamma' - \delta\alpha' + 2\delta\beta') \\
 & \quad = -\epsilon(1-a)\phi'\delta\phi' - e^{2\alpha}(2\delta\alpha W + W_\phi\delta\phi),
 \end{aligned}$$

$$\begin{aligned}
 & -e^{2(\alpha-\gamma)}\delta\ddot{\alpha} + \delta\gamma'' + 2\delta\beta'' - \delta\alpha'(\gamma' + 2\beta') - \alpha'(\delta\gamma' + 2\delta\beta') + 2\delta\gamma\gamma' + 4\beta'\delta\beta' \\
 & \quad = -\epsilon(3-a)\phi'\delta\phi' - e^{2\alpha}(2\delta\alpha W + W_\phi\delta\phi),
 \end{aligned}$$

$$\begin{aligned}
 & -e^{2(\alpha-\gamma)}\delta\ddot{\beta} + \delta\beta'' + \delta\beta'(\gamma' - \alpha' + 2\beta') + \beta'(\delta\gamma' - \delta\alpha' + 2\delta\beta') - 2e^{2(\alpha-\beta)}(\delta\alpha - \delta\beta) \\
 & \quad = -\epsilon(1-a)\phi'\delta\phi' - e^{2\alpha}(2\delta\alpha W + W_\phi\delta\phi),
 \end{aligned}$$

$$\delta\dot{\beta}' + (\beta' - \gamma')\delta\dot{\beta} - \beta'\delta\dot{\alpha} = -\frac{\epsilon}{2}\phi'\delta\dot{\phi},$$

$$\begin{aligned}
 & -e^{2(\alpha-\gamma)}\delta\ddot{\phi} + a\delta\phi'' + [\gamma' - a\alpha' + 2\beta']\delta\phi' + [\delta\gamma' - a\delta\alpha' + 2\delta\beta']\phi' \\
 & \quad = \epsilon e^{2\alpha}(2\delta\alpha W_\phi + W_{\phi\phi}\delta\phi),
 \end{aligned}$$

where $W(\phi) = (3 - 2a)V(\phi)$.

From the previous equations we can find the expression

$$\delta\gamma' = \eta\delta\phi' - \eta'\delta\phi,$$

with $\eta = \epsilon\phi'/(2\beta')$.

However, $\delta\gamma'$ may be alternatively calculated, which gives

$$2\beta'\delta\gamma' = \epsilon a\phi'\delta\phi' + e^{2\alpha}\left[2(e^{-2\beta} - W)\delta\alpha - W_\phi\delta\phi\right],$$

The comparison between the two expressions for $\delta\gamma'$ gives us

$$(1 - a)\phi'\delta\phi' = (1 - a)[\phi'' - \alpha'\phi' + \eta\phi'^2]\delta\phi$$

and its integration with respect to u give us

$$\delta\phi = F(t)\phi'e^{\alpha+Q(u)}, \quad Q(u) = \int \eta(u)\phi'(u)du,$$

Rastall perturbation: $b = -1$, $V = 0$

For this solution, we have the following master equation

$$e^{-4\gamma}\delta\ddot{\phi} + \delta\phi'' + \frac{2}{x}\delta\phi' = 0.$$

We obtain two expressions for $\delta\gamma'$

$$\delta\gamma' = \sqrt{3/2}\delta\phi' = -\sqrt{3/2}\delta\phi' - \sqrt{6}e^{2\alpha-2\beta}\delta\phi.$$

Which is integrated in relation to x , we obtemos

$$\delta\phi = F(t)\psi(x), \quad \psi(x) = (K - (C/\sqrt{6})x^4)^{-1/2}.$$

By separation of variables, we also obtain

$$-\frac{\ddot{F}}{F} = \frac{\psi''}{\psi} + \frac{2}{x}\frac{\psi'}{\psi} = \text{constant},$$

which is not satisfied by background solution.

- K-essence

Both solutions, for $n = 1/3$ and $n = 1/2$, are unstable.

On the instability of some k-essence space-times

(Bronnikov et al., 2019; [arXiv:1908.09126](#))

- Rastall gravity

Stability analysis in the Rastall case is inconsistent: linear time-dependent spherically symmetric perturbations simply do not exist.

Rastall's theory of gravity: Spherically symmetric solutions and the stability problem

(Bronnikov et al., 2020; [arXiv:2007.01945](#))

Obrigado!